

Generative strategies to empower physics-based wave propagation with deep learning. Applications to earthquake engineering.

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Grants

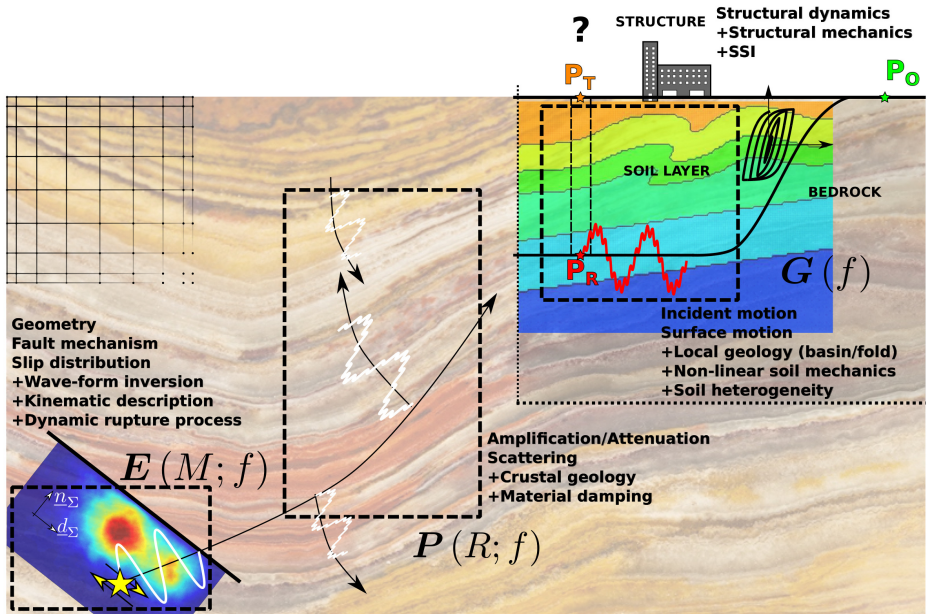
NVIDIA 40000 GPU hours

IDRIS 110000 GPU hours

CINES 6700000 CPU hours

General context

SOURCE-TO-STRUCTURE ANALYSIS



What we do

SEM3D [1, 6] : an open-source HPC code for earthquake simulation

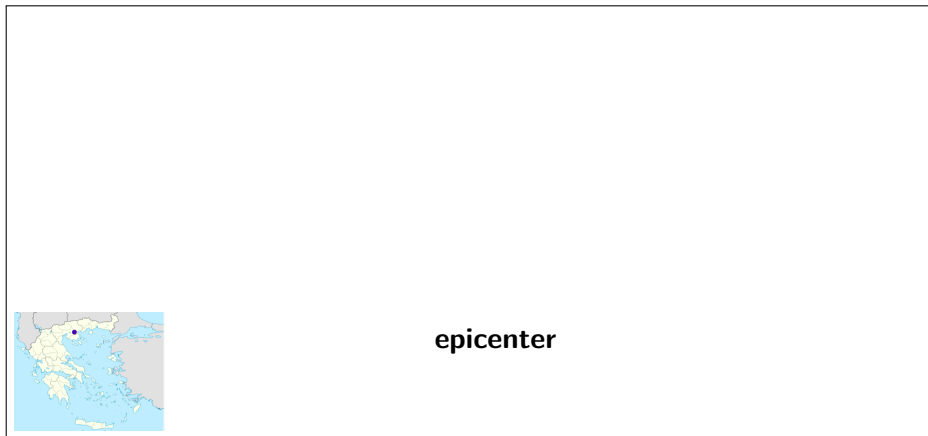


Figure – Earthquake simulation at the Volvi basin (Greece) [2]. 13 km × 11 km wide area. ≈7 h wall-time - 216 MPI cores @Ruche (≈62 days CPU time)

The earthquake engine : SEM3D

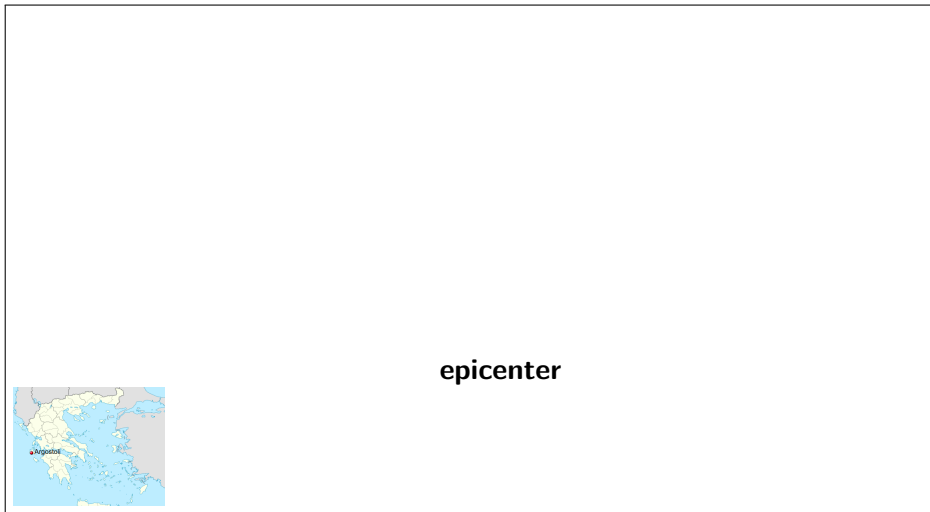


Figure – $1.3 \cdot 10^{10}$ dofs 0-10 Hz Argostoli earthquake [6]. ≈ 20 h wall-time - 4000 MPI cores @Occigen (3333 days CPU time)

Earthquake simulation at several nuclear facilities

Dynamic SSI simulation of Unit 7 reactor building at Kashiwazaki-Kariwa NPP (Japan). Weak-coupling SEM3D/code_aster [3].

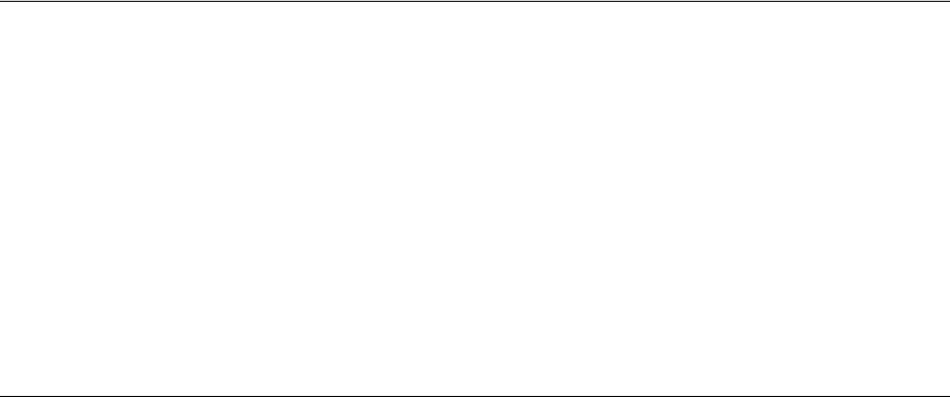
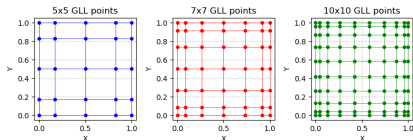
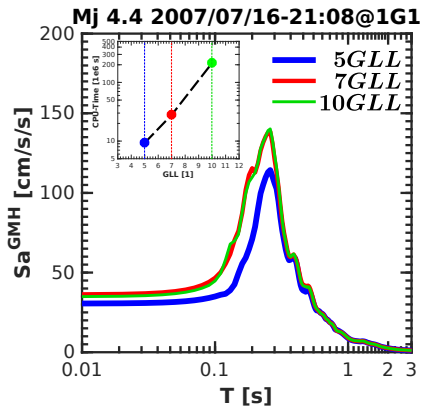


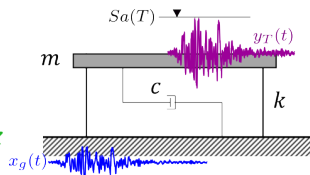
Figure – Courtesy of M. Korres EDF R&D.

Test numerical accuracy

All time-histories filtered at 7 Hz



$$f_{max}^{5GLL} \approx 3.5\text{Hz}; f_{max}^{7GLL} \approx 5.0\text{Hz}; f_{max}^{10GLL} \approx 7.0\text{Hz}$$



Spinning-off : EASYRISK in a nutshell




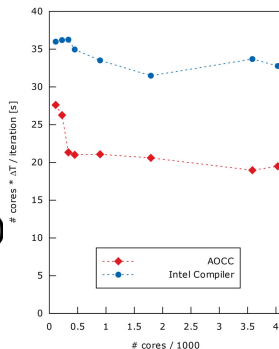
SATT Paris-Saclay POC'UP 2023 funding

EASYRISK : an user-friendly platform for physics-based earthquake simulation and risk assessment at regional scale

- ▶ Pre-/Post-processing platform for SEM3D
- ▶ Use Machine Learning and BigData tools
- ▶ End-to-end seismic risk assessment solutions
- ▶ Simulate blind earthquake scenarios

Make seismic **RISK** estimation **EASY** (on cloud)

Checkout [blog article](#) in collaboration with 

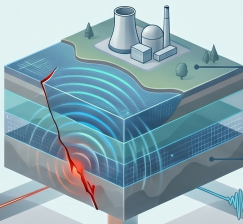


Earthquake Digital Twins: The Power and Open Challenges of Physics-Based Simulation

WHAT PHYSICS-BASED SIMULATIONS (SEM3D) CAN DO

Source-to-Structure Analysis

SEM3D models the complete physics of an earthquake, including active fault rupture, wave scattering through crustal geology, and non-linear soil mechanics at the surface.



Billions of Degrees of Freedom

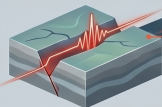
State-of-the-art simulations, such as the Argostoli earthquake model, involve up to 15 billion degrees of freedom to capture regional wave propagation.



Critical Infrastructure Protection

These simulations are used to assess the safety of nuclear facilities by modeling how reactors react to dynamic seismic loads in 3D.

THE REALITY OF RETRO-ENGINEERING PAST EARTHQUAKES



The 0-2 Hz Frequency Bottleneck

Current accurate retro-engineering is largely restricted to a limited frequency range (0-2 Hz or 0-18 Hz), which misses higher-frequency "vibrational fingerprints" critical for structural design.



Hard but Not Impossible

Reconstructing past earthquakes is achievable through high-fidelity digital twins, but it is limited by significant technical constraints.



High Inherent Uncertainty

Models suffer from poor resolution compared to the intricacy of the Earth's crust, and accurate geological measurements (like shear wave velocity) are often unavailable.



THE OPEN CHALLENGES: LOOKING TO THE FUTURE



The Problem of Multiple Geological Models

Researchers must determine how to synthesize results and manage high variability when multiple conflicting geological interpretations of the same site exist.



Missing Observations

A major hurdle is how to build reliable physics-based models for regions where no historical earthquake recordings or borehole data are available.



Unlocking PSHA via Reduced Costs

To enable Probabilistic Seismic Hazard Assessment (PSHA), we must reduce the "prohibitive" computational cost of running thousands of scenario simulations.



Real-Time Multi-Scenario Simulation

The ultimate goal is to move from months of supercomputing to real-time simulation of multiple scenarios, likely through AI surrogate models like Diffusion Transformers (DITD).

NotebookLM

M_W 4.9 Le Teil earthquake (2019) [4]

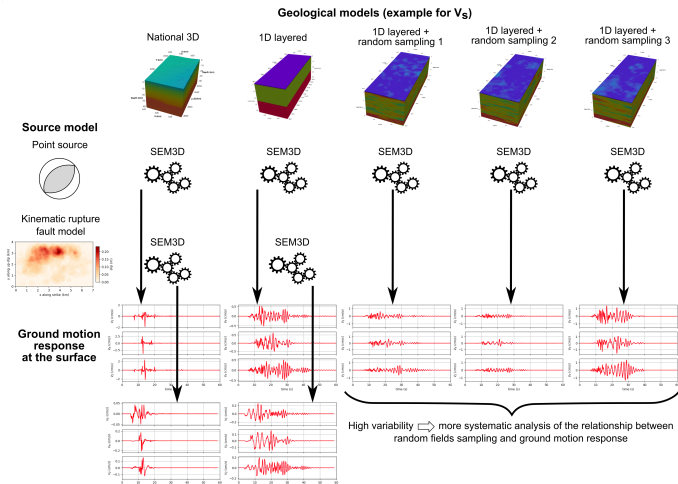
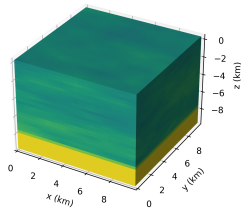
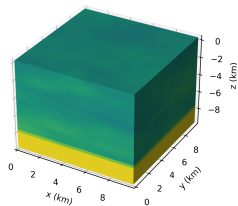


Figure – 80 km \times 92 km \times 79 km - 18.3 million hexahedra. 60 s of simulated signal
61440h CPU time - 2048 MPI cores AMD Milan @2.45GHz (AVX2) on TGCC.

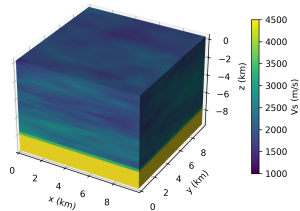
Input



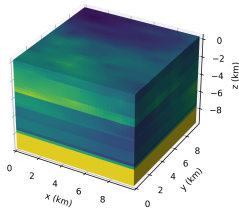
PCA 1024 components (RMSE=35.9m/s)



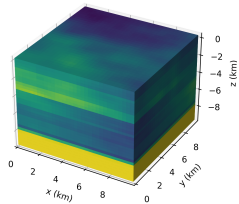
3D UNet (RMSE=761.2m/s)



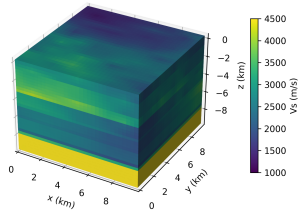
Input



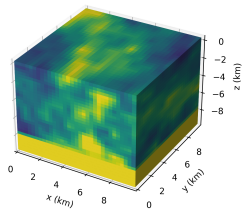
PCA 1024 components (RMSE=45.2m/s)



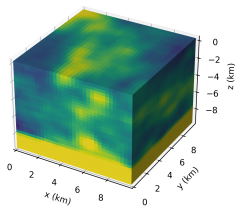
3D UNet (RMSE=70.6m/s)



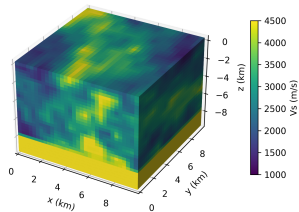
Input



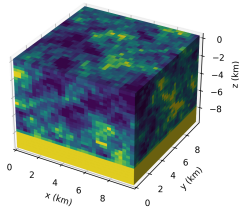
PCA 1024 components (RMSE=133.3m/s)



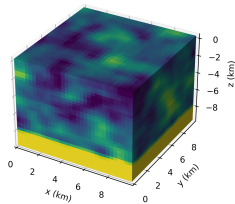
3D UNet (RMSE=136.7m/s)



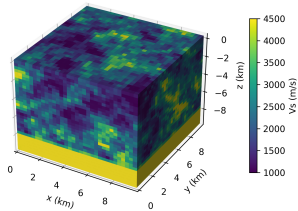
Input



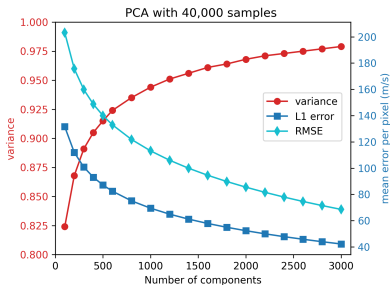
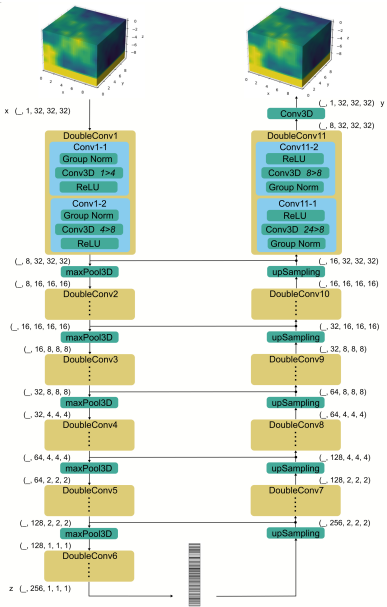
PCA 1024 components (RMSE=405.9m/s)



3D UNet (RMSE=47.6m/s)



Dimensionality reduction [4]

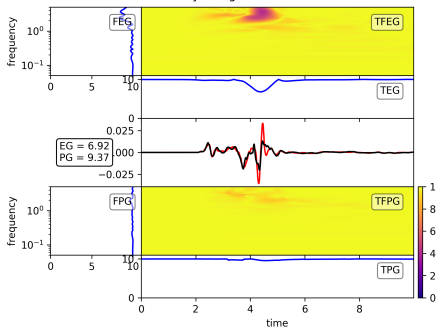


- ▶ PCA : 500 eigen-geologies explain 91.% of the variance (1000 eigen-geologies : 94.4%, 3000 eigen-geologies : 97.9%)
- ▶ Similar results with a 3D UNet of 4000000 weights trained over 10000 epochs (4 Nvidia Ampere A100 GPUs took 9.5 hours)

Dimensionality reduction [4]

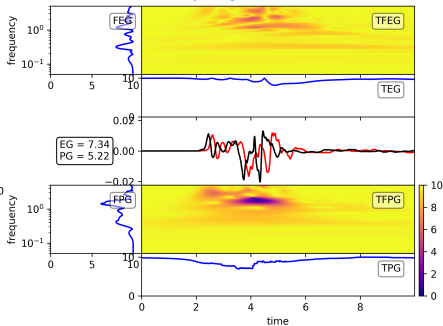
PCA 1024 components

Velocity along the N-S axis



3D UNet

Velocity along the N-S axis

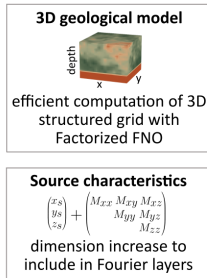


Machine learning opportunities to conduct high-fidelity earthquake simulations in multi-scale heterogeneous geology [4]

- ▶ Can we entirely replace simulations with AI? (neural operators might...see you tomorrow)
- ▶ Can we use AI to enlarge the frequency range?

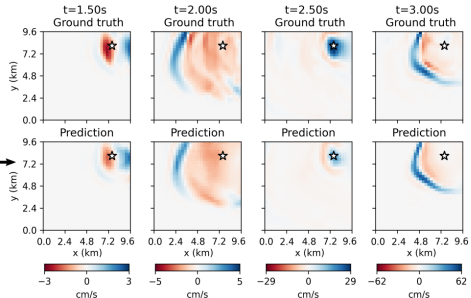
Earthquake neural simulator

Multiple-Input FNO for source-dependent 3D elastodynamics



MIFNO →

Surface velocity wavefields

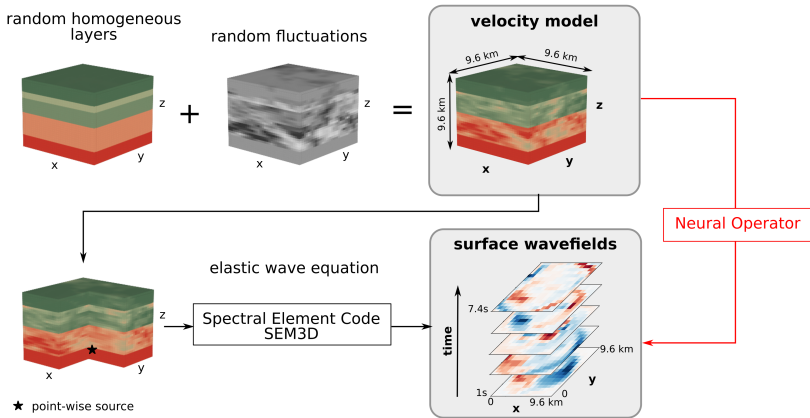


Generalization to:

- ✓ Not autoregressive \Rightarrow long-term stability
- ✓ Excellent wave arrival times
- ✓ Good to excellent amplitudes
- ✓ Out-of-distribution sources
- ✓ Overthrust geologies
- ✓ Higher resolution

Neural operators

HEMEW-3D ([link](#)) : 30000+ SEM3D simulations at 5 Hz (open-source !)



Sought non-linear mapping $\mathcal{G} : L^2(\Omega) \ni c_S \mapsto \mathbf{v}|_{\partial\Omega_n} \in \mathcal{S}^t$

$$\mathcal{S}^t := \left\{ \mathbf{v} : \Omega \times [0, T] \rightarrow \mathbb{R}^3 : \mathbf{v} \in H^1(\Omega)^3 \forall t \in [0, T] \right\}$$

Neural operator : $\mathcal{G}_\theta(c_S)(\mathbf{x}) = \int_\Omega \mathcal{K}_\theta(\mathbf{x}, \mathbf{y}, c_S(\mathbf{x}), c_S(\mathbf{y})) c_S(\mathbf{y}) d\mathbf{y}$

Multiple-Input Fourier Neural Operator (MIFNO)

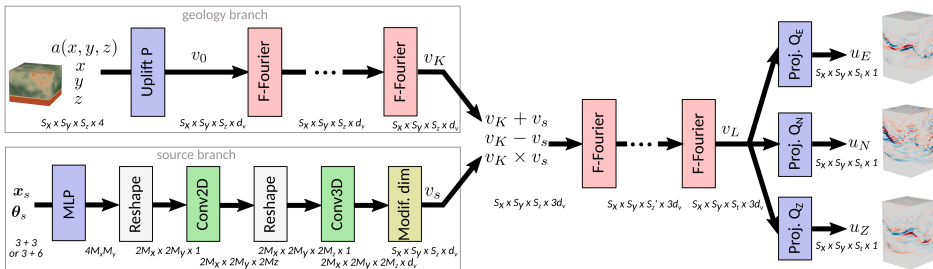


Figure — The MIFNO is made of a *geology branch* that encodes the geology with factorized Fourier (F-Fourier) layers, and a *source branch* that transforms the vector of source parameters (x_s, θ_s) into a 4D variable v_s matching the dimensions of the *geology branch* output v_K . Outputs of each branch are concatenated after elementary mathematical operations and the remaining factorized Fourier layers are applied. Uplift P and projection Q_E, Q_N, Q_Z blocks are the same as in the F-FNO.

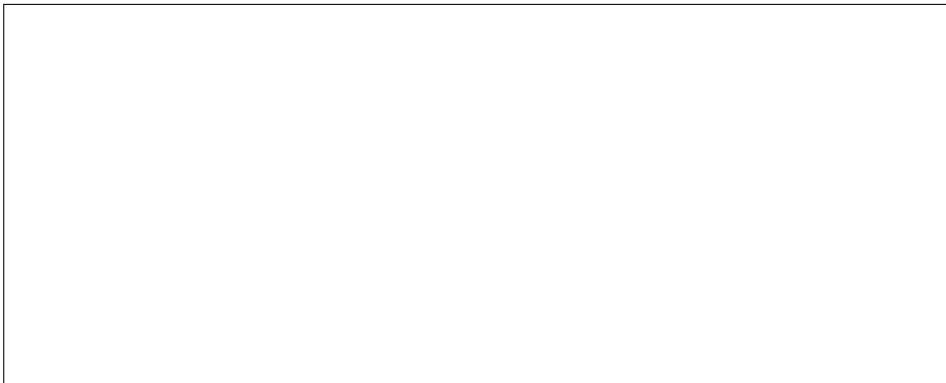
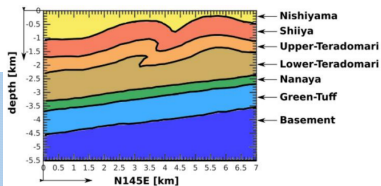


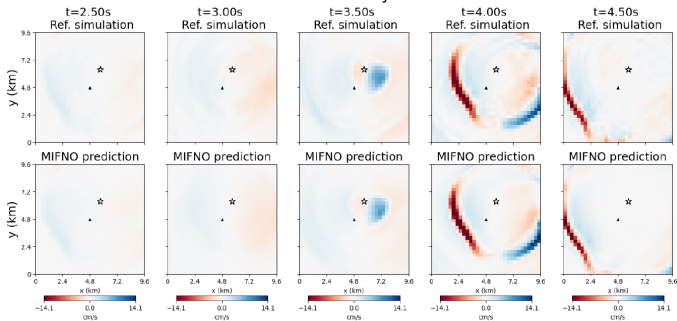
Figure – Source-to-site finite-fault simulation with MIFNO.

1. choose randomly one geology in the Le Teil database
2. for each point on the fault plane, predict the surface wavefields with the fine-tuned MIFNO
3. deconvolve all surface wavefields by the spice bench source time function
4. convolve each surface wavefield by the source time function corresponding to its associated point source
5. sum all surface wavefields

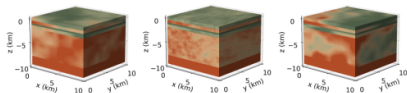
Niigata earthquake 2007



East-West velocity field

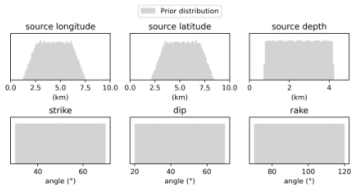


Zero-shot MIFNO

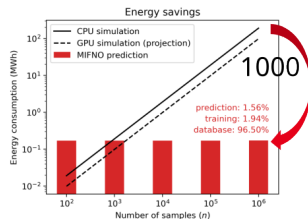


✗ Some inputs may be incompatible with reality

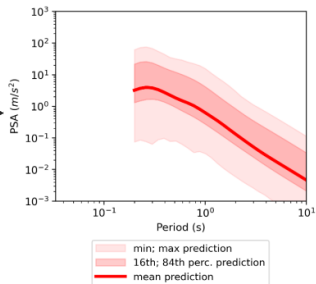
prior distribution



Fast predictions



PSA distribution

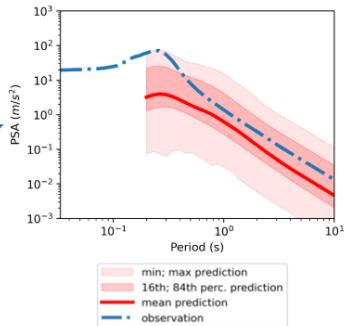
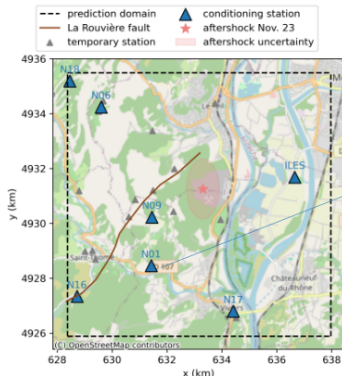


✗ PSA range is very large

Best-matching geologies

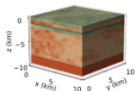
A small-magnitude event with PSA has been observed

Le Teil aftershock November 23



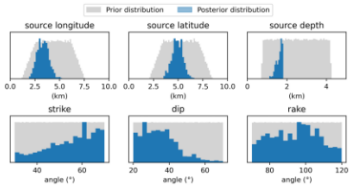
Best-matching geologies (cont'd)

posterior distribution



✓ We have selected a set of admissible geologies

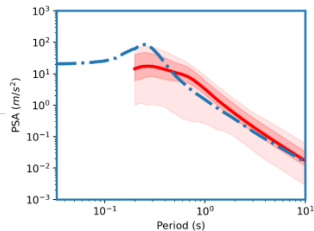
posterior distribution



✓ We are able to constrain the source parameters

Select inputs that satisfy the observation

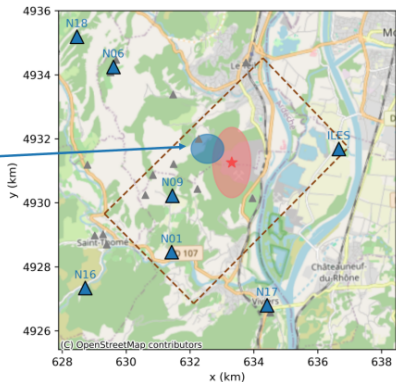
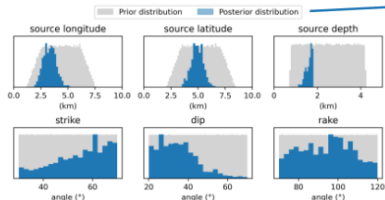
PSA distribution



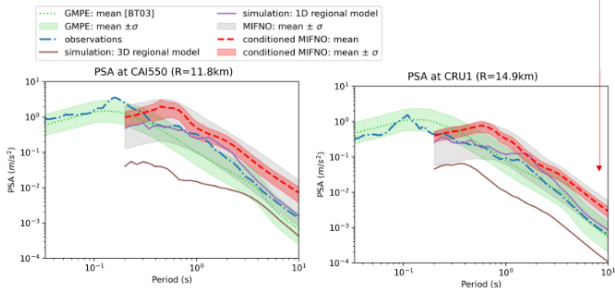
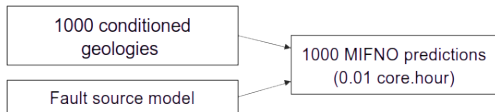
Best-matching sources

« Free » Bayesian solution to the inverse problem coherent with conventional inversion techniques

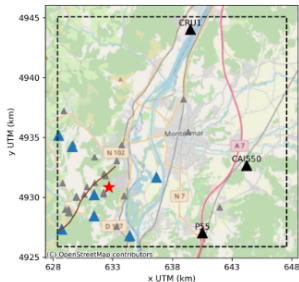
posterior distribution



Data-tuned MIFNO



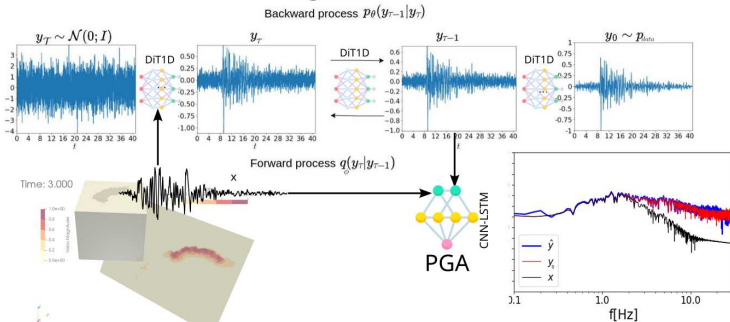
- ▲ temporary stations
- ▲ conditioning stations
- ▲ permanent stations
- prediction domain (19.2km x 19.2km)
- La Rouvière fault
- ★ mainsshock



- ✓ Conditioned distributions
- ✓ For a strong event
- ✓ On permanent stations

Earthquake neural simulator

Conditioned denoising diffusion on seismic records



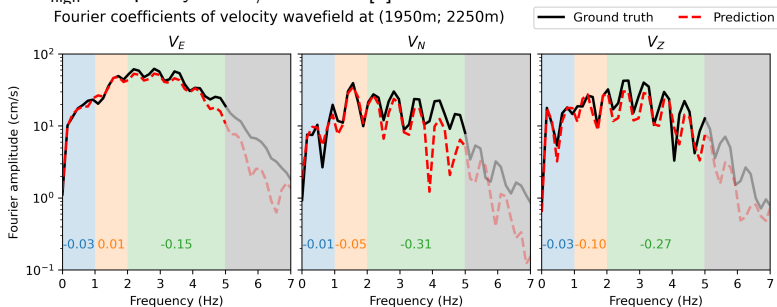
Conditioning PBS

The “plague” of spectral bias

Model	rRMSE	rFFT _{low}	rFFT _{mid}	rFFT _{high}
FNO	0.23 ± 0.088	-0.17 ± 0.18	-0.30 ± 0.24	-0.44 ± 0.30
U-NO	0.22 ± 0.088	-0.16 ± 0.16	-0.28 ± 0.22	-0.44 ± 0.28
G-FNO	0.26 ± 0.073	-0.53 ± 0.12	-0.76 ± 0.11	-0.94 ± 0.04
F-FNO $N^{(\ell)} = 4$	0.23 ± 0.081	-0.29 ± 0.16	-0.42 ± 0.21	-0.55 ± 0.26
F-FNO $N^{(\ell)} = 8$	0.21 ± 0.083	-0.15 ± 0.14	-0.26 ± 0.19	-0.37 ± 0.26
F-FNO $N^{(\ell)} = 28$	0.17 ± 0.081	-0.08 ± 0.10	-0.14 ± 0.15	-0.23 ± 0.22

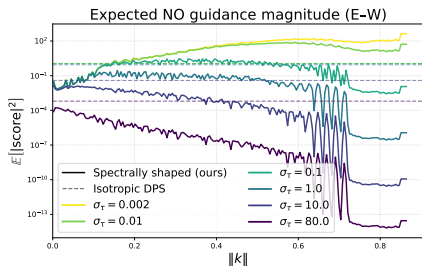
Table – Mean and standard deviation of relative biases for 1000 validation samples. rFFT_{low}, rFFT_{mid}, rFFT_{high} : frequency biases/deviations [5].

Fourier coefficients of velocity wavefield at (1950m; 2250m)



FreqNO-DPS : fuse a diffusion prior, sensors, and a frozen NO

$$\nabla_{\mathbf{u}_\tau} \log p_\tau(\mathbf{u}_\tau \mid \mathbf{y}, \mathbf{u}_{\text{NO}}) \approx \underbrace{\frac{\bar{\mathbf{u}}_0 - \mathbf{u}_\tau}{\sigma_\tau^2}}_{\text{prior (Tweedie)}} + \underbrace{\nabla_{\mathbf{u}_\tau} \log p(\mathbf{y} \mid \mathbf{u}_\tau)}_{\text{sensor anchor}} + \underbrace{\nabla_{\mathbf{u}_\tau} \log p(\mathbf{u}_{\text{NO}} \mid \mathbf{u}_\tau)}_{\text{NO guidance}}$$



Naive (isotropic) NO guidance reimports the surrogate's bias.

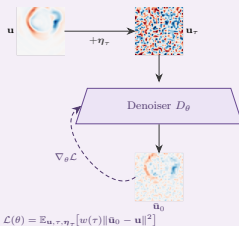
(1)

Observations & calibration*precomputed, one-time*

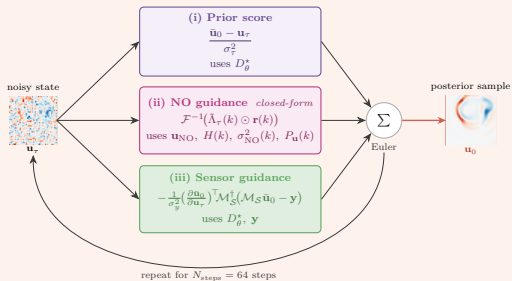
Spectral calibration
 $H(k), \sigma_{NO}^2(k), P_u(k)$
 from paired (u, u_{NO})



(2)

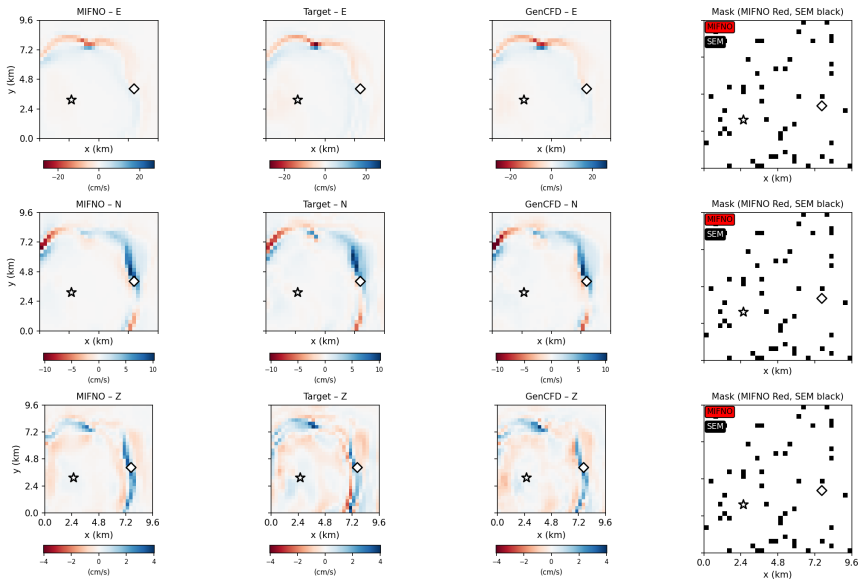
Denoiser training*on SEM ground truth wavefield u* 

(3)

Inference: posterior sampling*one reverse-diffusion step on $\nabla_{u_\tau} \log p_\tau(u_\tau | y, u_{NO})$* 

Spoiler alert !

Velocity Field Comparison at PGV (target-based) — $t \approx 2.840s$



The key idea : spectrally shaped NO guidance

Per-mode error model, calibrated from paired data :

$$\mathcal{F}(\mathbf{u}_{\text{NO}})(\mathbf{k})_c = H(\mathbf{k}) \mathcal{F}(\mathbf{u})(\mathbf{k})_c + \hat{\eta}_{\text{NO}}(\mathbf{k}), \quad \mathbb{E}|\hat{\eta}_{\text{NO}}|^2 = \sigma_{\text{NO}}^2(\mathbf{k}).$$

Two-step LMMSE, with $X = \mathcal{F}(\mathbf{u})_c$, $Y = \mathcal{F}(\mathbf{u}_\tau)_c$, $Z = \mathcal{F}(\mathbf{u}_{\text{NO}})_c$:

Step 1 — diffusion channel : estimate clean X from the noisy state Y

Wiener filter α (shrinks Y toward the prior mean) and its error σ_L^2 :

$$\alpha(\mathbf{k}) = \frac{P_{\mathbf{u}}(\mathbf{k})}{\sigma_\tau^2 + P_{\mathbf{u}}(\mathbf{k})}, \quad \sigma_L^2(\mathbf{k}) = \sigma_\tau^2 \alpha(\mathbf{k}) \leq \min(P_{\mathbf{u}}, \sigma_\tau^2).$$

Step 2 — NO channel : refine with the surrogate Z

Innovation $r(\mathbf{k}) = Z - H(\mathbf{k}) \alpha(\mathbf{k}) Y$ has variance

$$\lambda_\tau(\mathbf{k}) = |H(\mathbf{k})|^2 \sigma_L^2(\mathbf{k}) + \sigma_{\text{NO}}^2(\mathbf{k}).$$

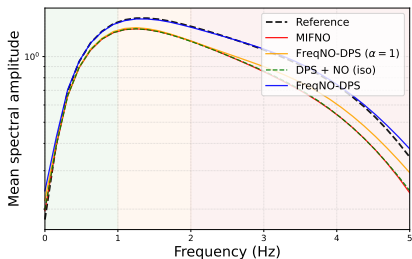
Result : near-zero spectral bias at 2–5% coverage

Closed-form score — two FFTs, **no denoiser backprop** :

$$\nabla_{\mathbf{u}_\tau} \log p(\mathbf{u}_{\text{NO}} | \mathbf{u}_\tau) = \mathcal{F}^{-1}(\tilde{\lambda}_\tau \odot r), \quad \tilde{\lambda}_\tau(\mathbf{k}) = \frac{2 H(\mathbf{k}) \alpha(\mathbf{k})}{\lambda_\tau(\mathbf{k})} \xrightarrow{\text{high } \|\mathbf{k}\|} 0.$$

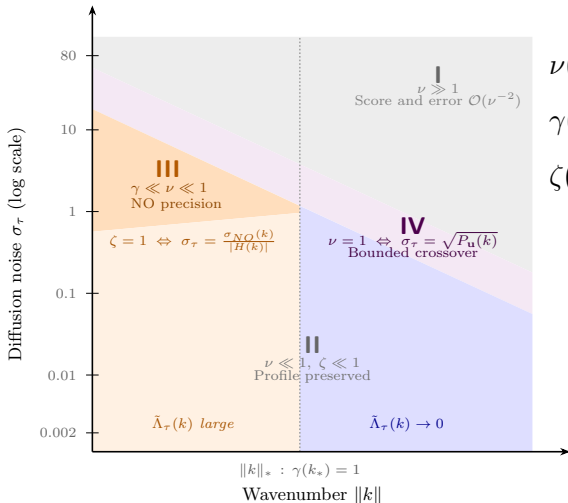
Method ($\rho = 5\%$)	rMAE \downarrow	rFFT _{high} $\rightarrow 0$
MIFNO (surrogate)	0.133	-0.239
DPS (sensors only)	0.113	-0.232
DPS + NO (iso)	0.099	-0.210
FreqNO-DPS	0.100	+0.002

Isotropic guidance improves pointwise error *but keeps the bias*. FreqNO-DPS reaches **near-zero bias at no pointwise cost** (+0.047 even at $\rho = 2\%$).



Ensemble-averaged spectrum, $\rho = 5\%$. Only FreqNO-DPS (blue) the reference across all bands.

Guiding regimes



$$\nu(\mathbf{k}, \tau) := \frac{\sigma_\tau^2}{P_u(\mathbf{k})}$$

$$\gamma(\mathbf{k}) := \frac{\sigma_{NO}^2(\mathbf{k})}{|H(\mathbf{k})|^2 P_u(\mathbf{k})}$$

$$\zeta(\mathbf{k}, \tau) := \frac{|H(\mathbf{k})|^2 \sigma_L^2(\mathbf{k})}{\sigma_{NO}^2(\mathbf{k})} = \frac{\nu}{\gamma(1+\nu)}$$

NO guidance ($\gamma < 1$) Sensors + prior ($\gamma > 1$) Prior only ($\nu \gg 1$)

Takeaways

- ▶ **Frequency-dependent calibration is essential, not merely beneficial** : it is what separates useful NO guidance from a faithful reimport of its bias.
- ▶ **Cheap and modular** : closed form, two FFTs per step, no denoiser backprop ; wraps *any* frozen surrogate.
- ▶ **Distribution-agnostic** : an exact score identity plus error bounds across the frequency–diffusion-time plane preserve the guidance's spectral shape regardless of distributional assumptions.
- ▶ **Lightweight prerequisites** : only paired surrogate/reference data for calibration, plus a coherence diagnostic to verify applicability to a new surrogate.

Applies broadly to FNO-family surrogates across wave propagation, fluids, and other PDE settings.

Thank you !
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Bibliography II

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